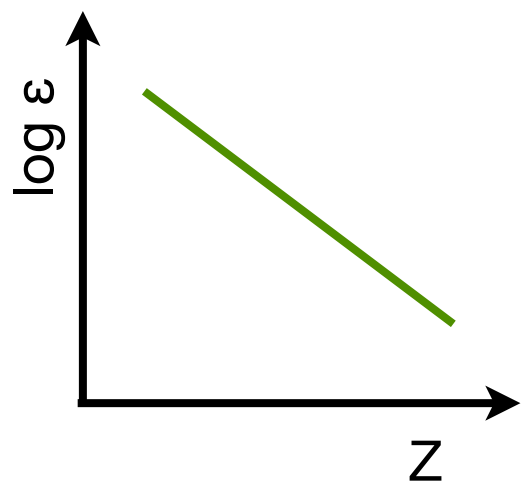


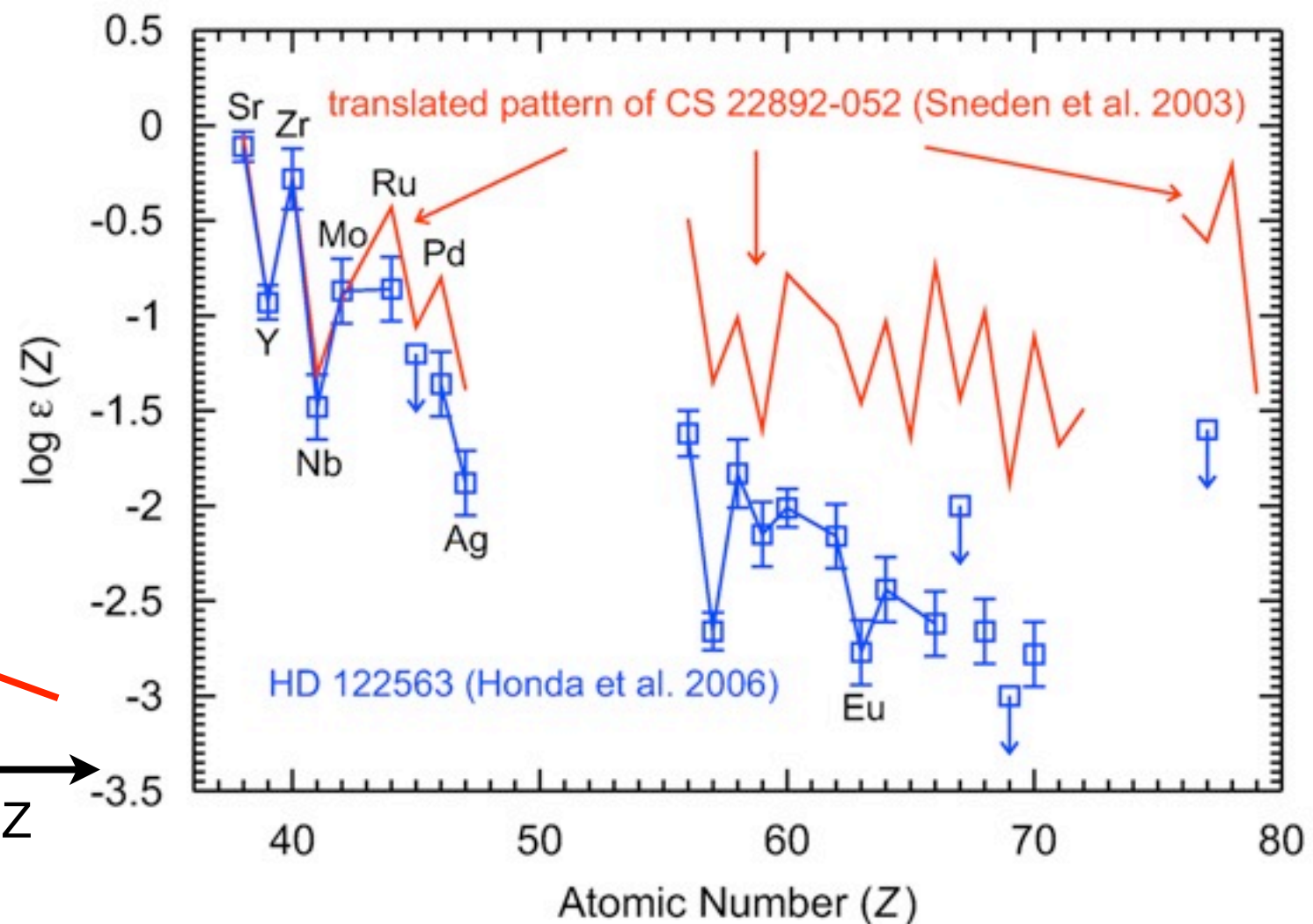
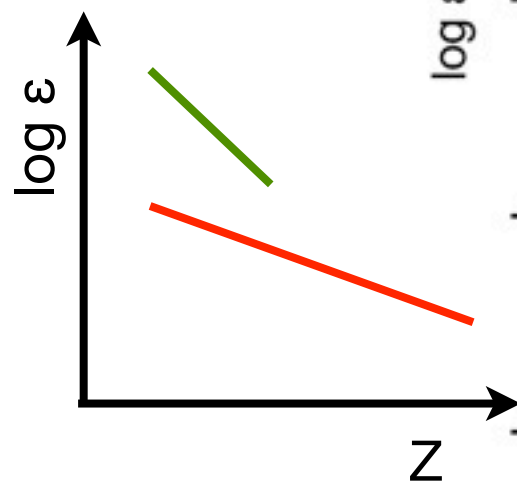
Lighter heavy elements: Sr - Ag

Ultra metal-poor stars with **high** and **low** enrichment of heavy r-process nuclei suggest: at least two components or sites (Qian & Wasserburg):

Are Honda-like stars the outcome of one nucleosynthesis event or the combination of several?



or



Travaglio et al. 2004: solar=r-process+s-process+LEPP
Montes et al. 2007: solar LEPP ~ UMP LEPP → unique

Nucleosynthesis components

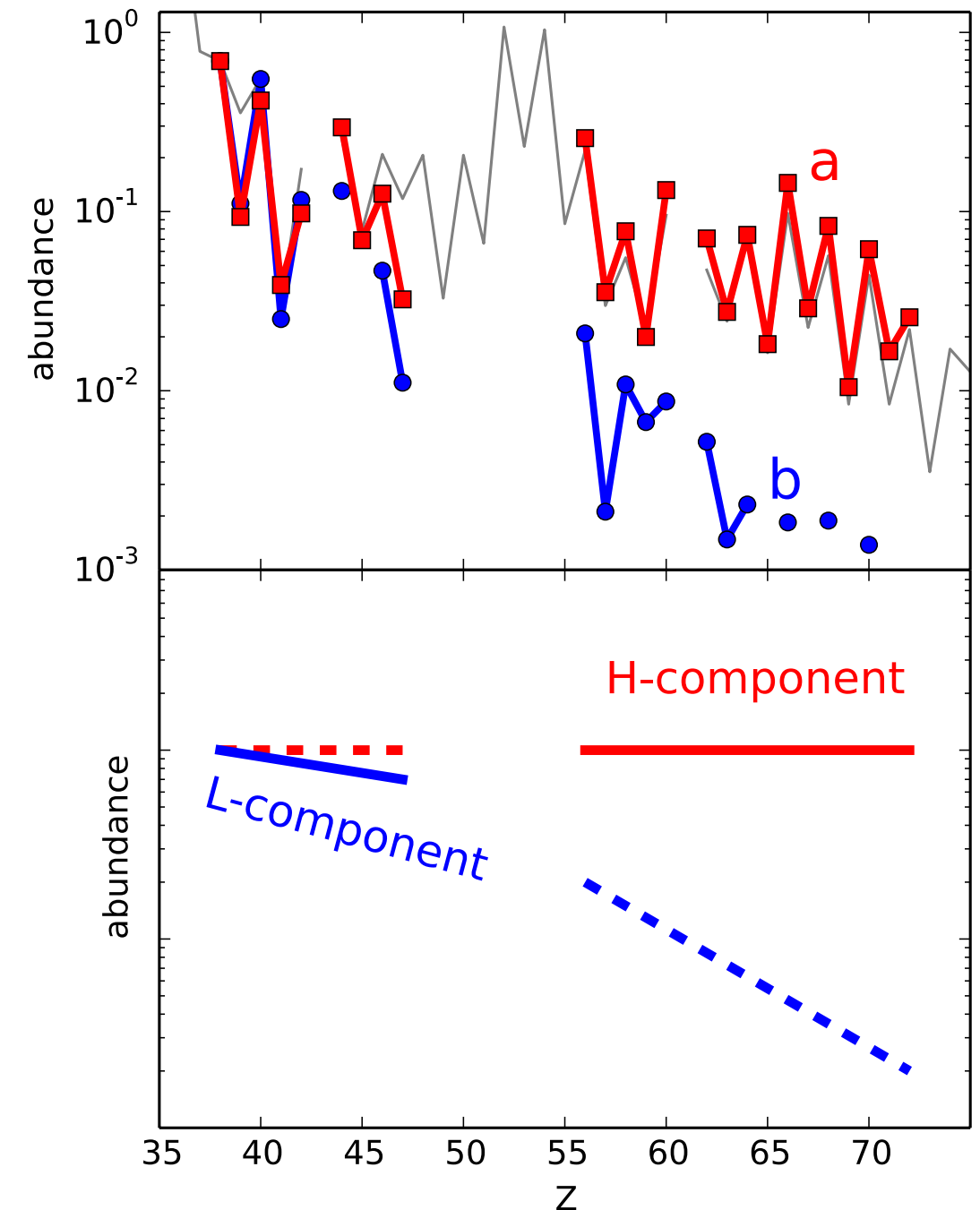
C.J. Hansen, Montes, Arcones 2014

L and H-components based on 3 methods:

M1: L-component = Honda star
H-component = Sneden star

M2: L-component = Honda - Sneden
H-component = Sneden

M3: iterative method (Li et al. 2013)
L-component = L - H
H-component = H - L



→ Component abundance pattern: Y_r and Y_L

Assumptions: Z range for components

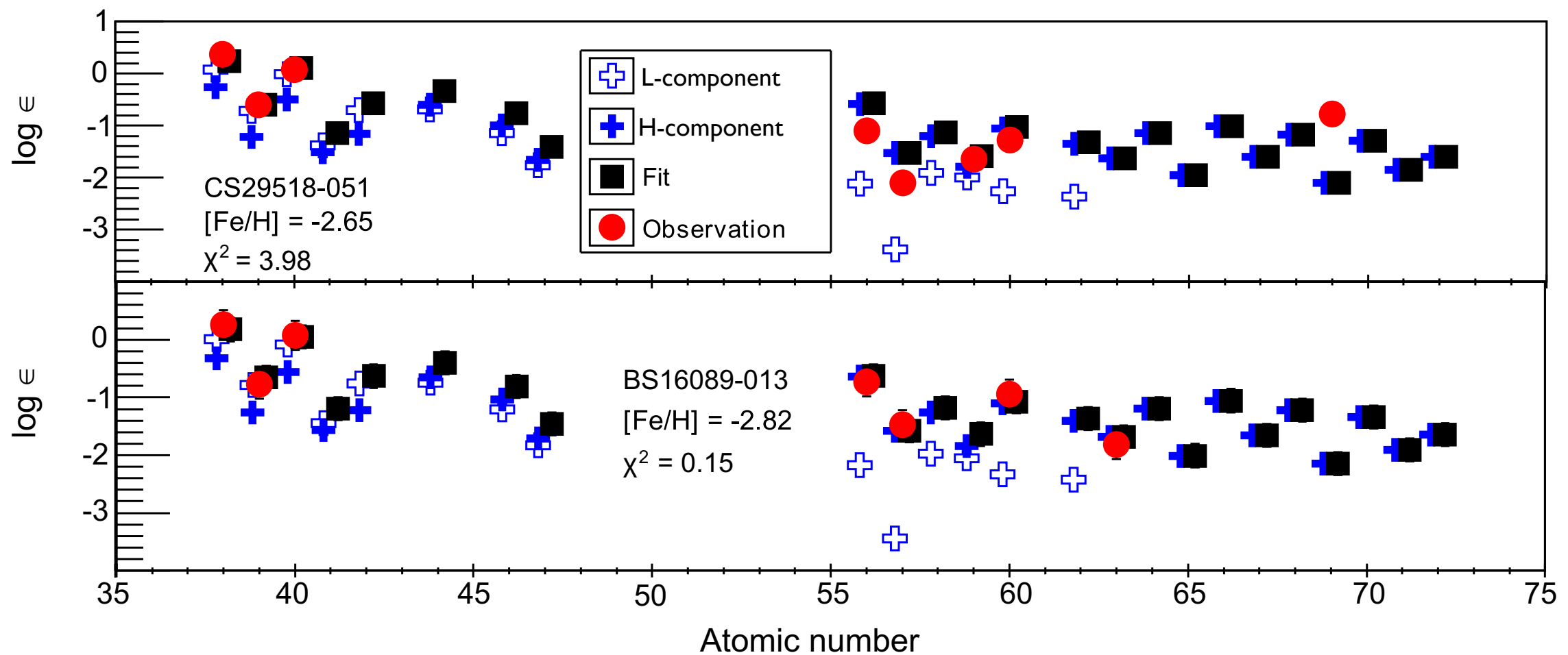
robust pattern with in uncertainties (0.32dex)

Abundance deconvolution

big sample of stars (Frebel et al. 2010)

remove s-process, carbon enhanced, and stars with internal mixing

fit abundance as combination of components: $Y_{\text{calc}}(Z) = (C_r Y_r(Z) + C_L Y_L(Z)) \cdot 10^{[\text{Fe}/\text{H}]}$



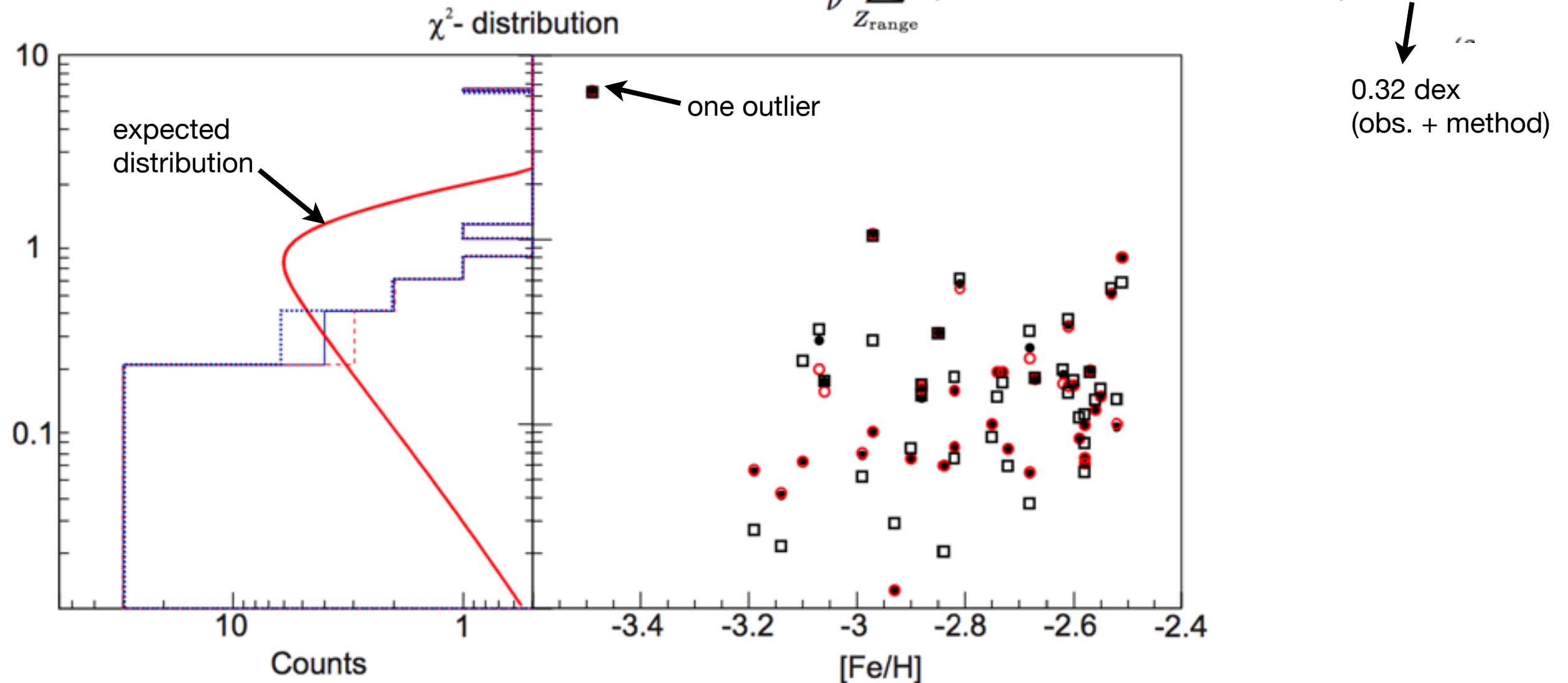
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$$\chi^2 = \frac{1}{\nu} \sum_{Z_{\text{range}}} (\log Y_{\text{observed}}(Z) - \log Y_{\text{calc}}(Z))^2 / \Delta(Z)^2,$$



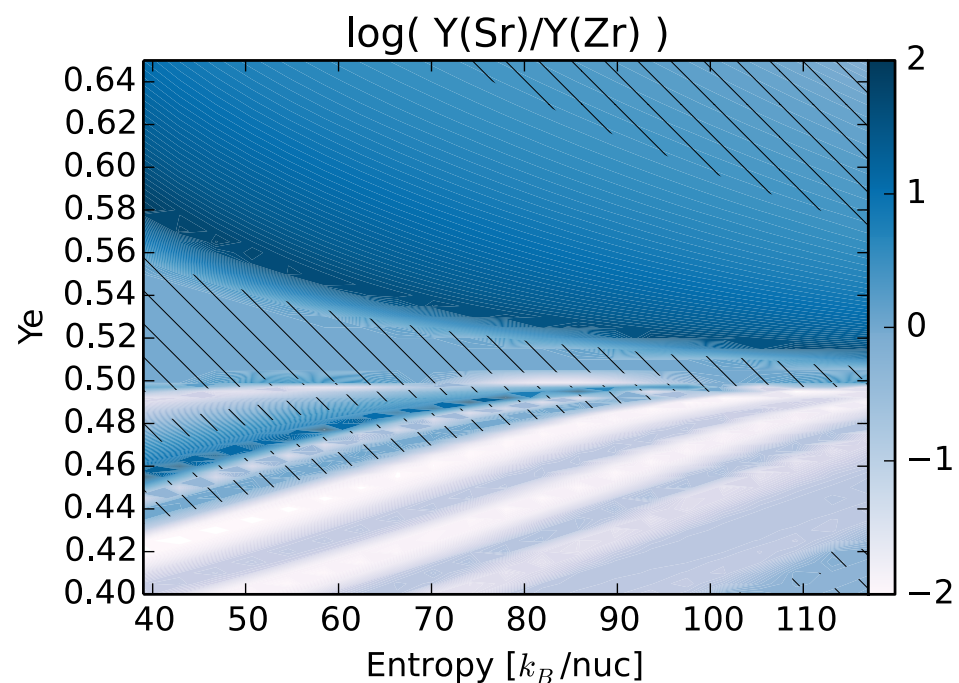
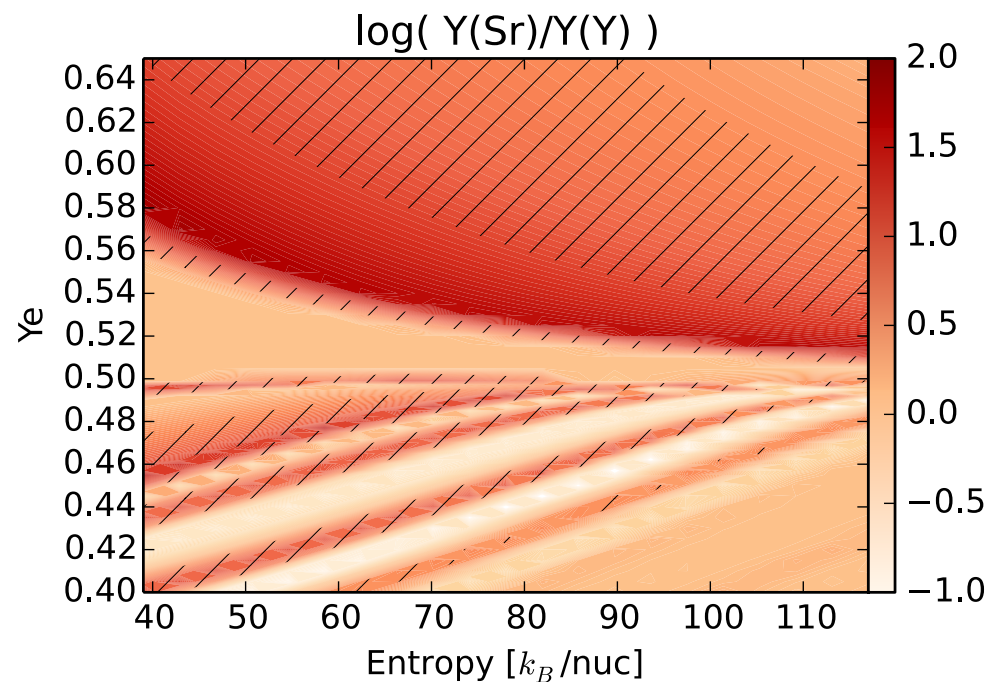
L-component: constraining conditions

L-component abundance ratios:

$$\text{Sr}/\text{Y} = 6.13 (/ /)$$

$$\text{Sr}/\text{Zr} = 1.22 (\\)$$

$$\text{Sr}/\text{Ag} = 48.2$$



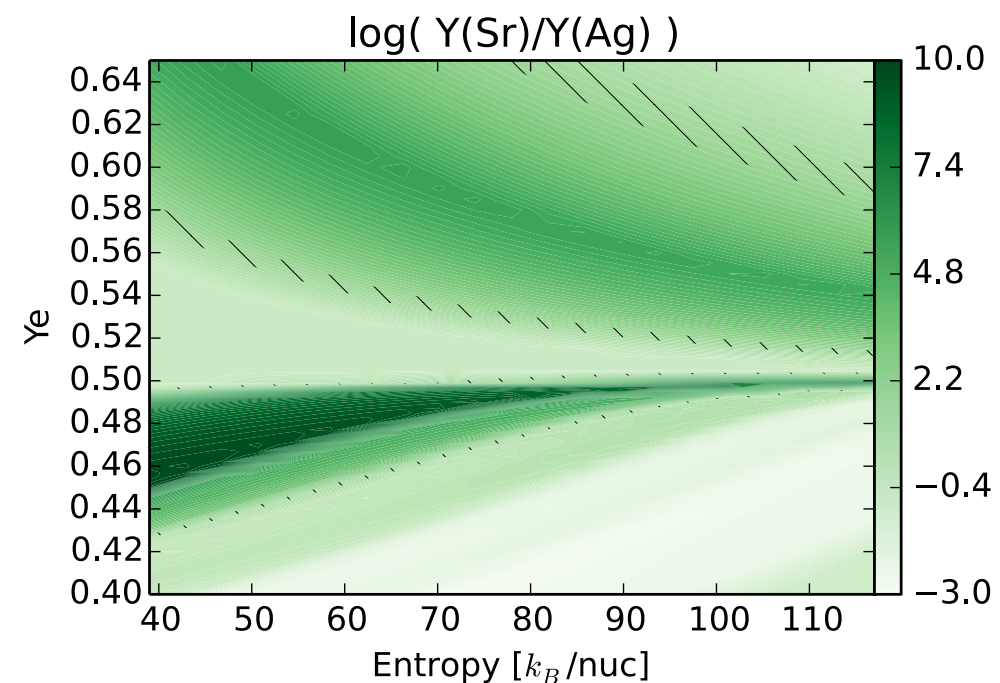
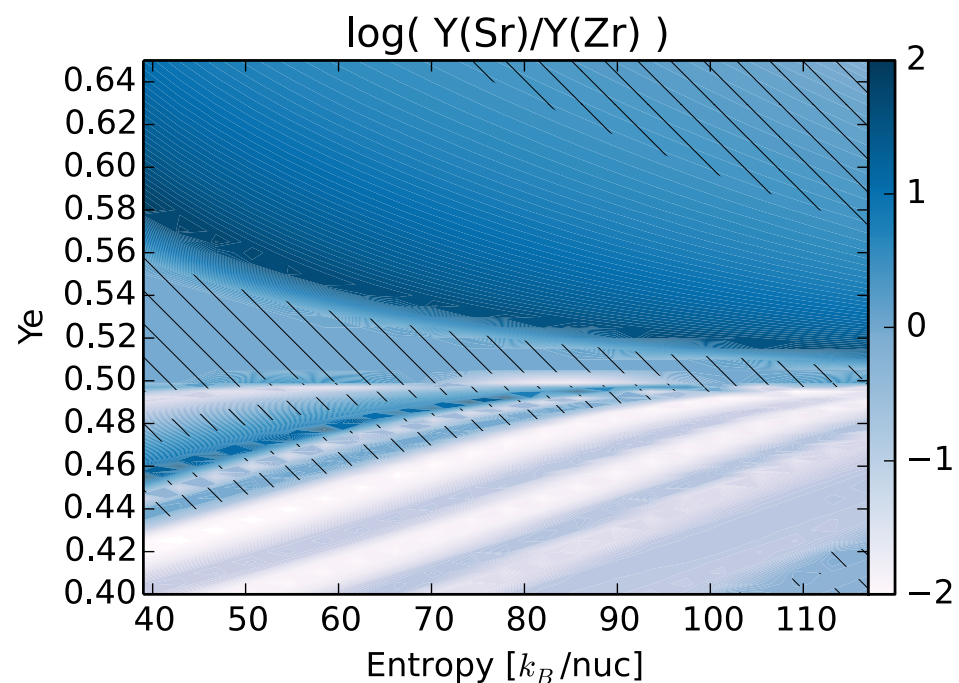
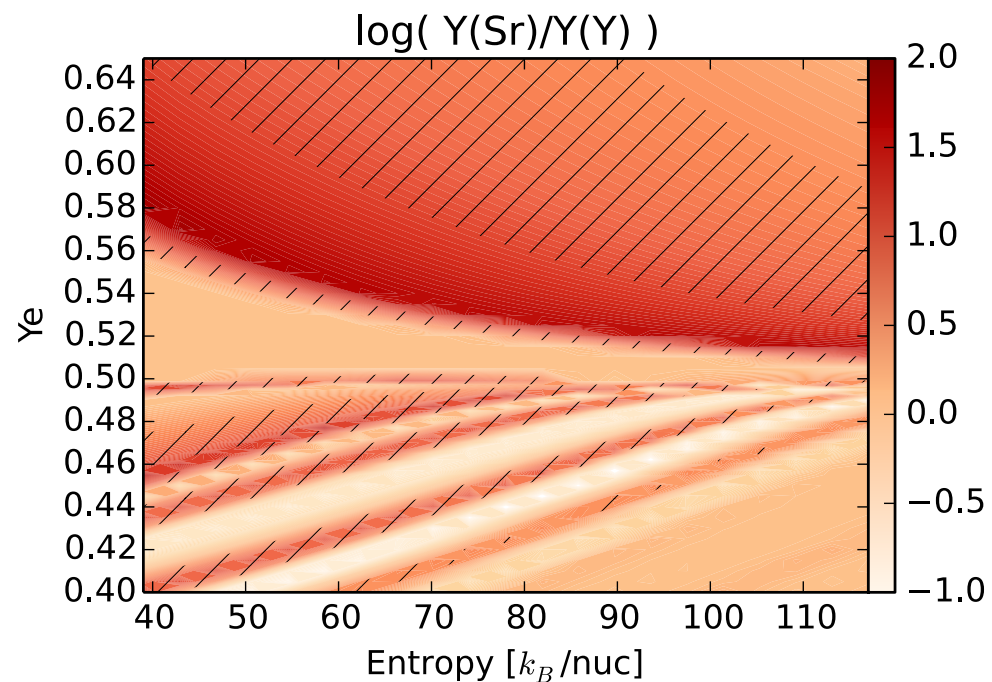
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